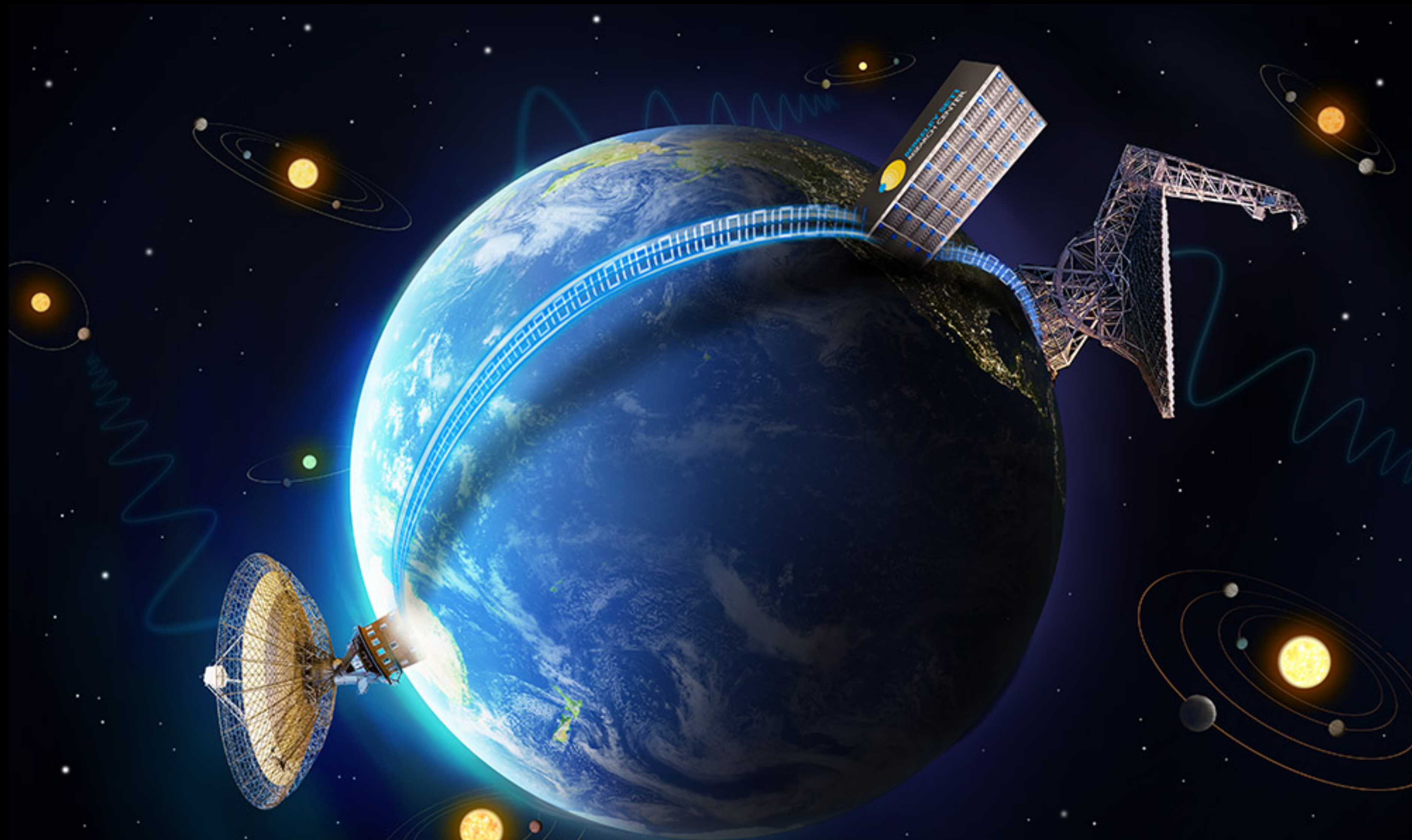


# BREAKTHROUGH LISTEN: FINDING LIFE BEYOND EARTH WITH HDF5



DR DANNY PRICE, ICRAR / UC BERKELEY  
EUROPEAN HDF USERS GROUP SUMMER 2021

**BREAKTHROUGH  
LISTEN**





03.2019

FLEEING CHAOS IN  
EL SALVADOR

CARNIVAL  
CELEBRATIONS

EXPLORING  
BORNEO'S CAVES

# NATIONAL GEOGRAPHIC

## WE ARE NOT ALONE

Scientists say there must be other life in the universe.  
Here's how they're searching for it.



*"Something great is  
around those stars."*  
SARA SEAGER,  
ASTROPHYSICIST



A deep space image showing a vast field of galaxies and stars against a black background. The galaxies are of various shapes and sizes, some appearing as bright, elongated structures, while others are smaller, distant points of light. The stars are scattered throughout, with some showing prominent diffraction spikes.

300,000,000,000

*Number of stars in the Milky Way (approx)*

>2,000,000,000,000

*Number of galaxies in observable Universe (approx)*



> 2000000000000000000000000

$$(2 \times 10^{23})$$

# *Estimated number of stars in the Universe*

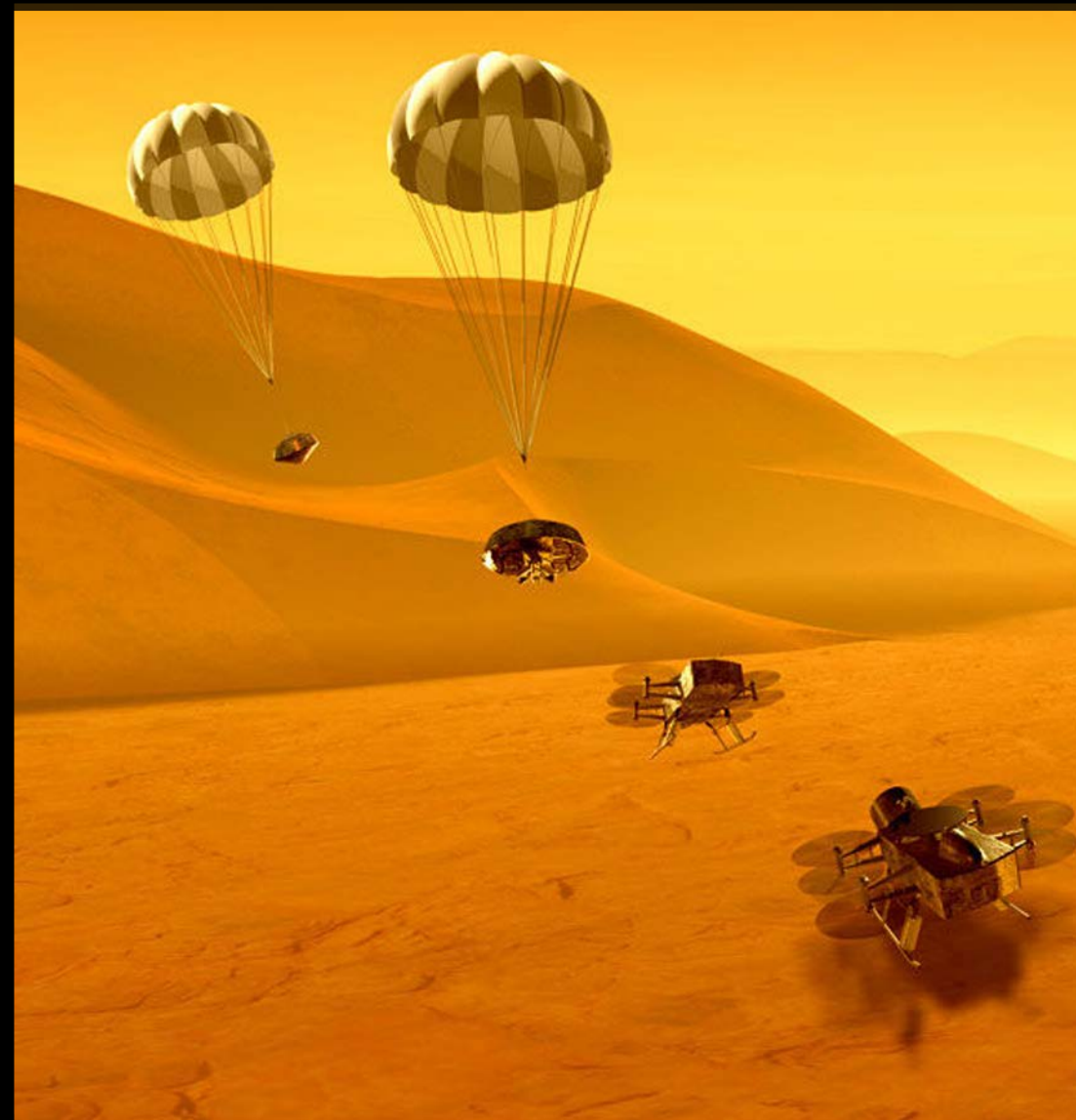


How do we find life beyond Earth?



# How do we find life beyond Earth?

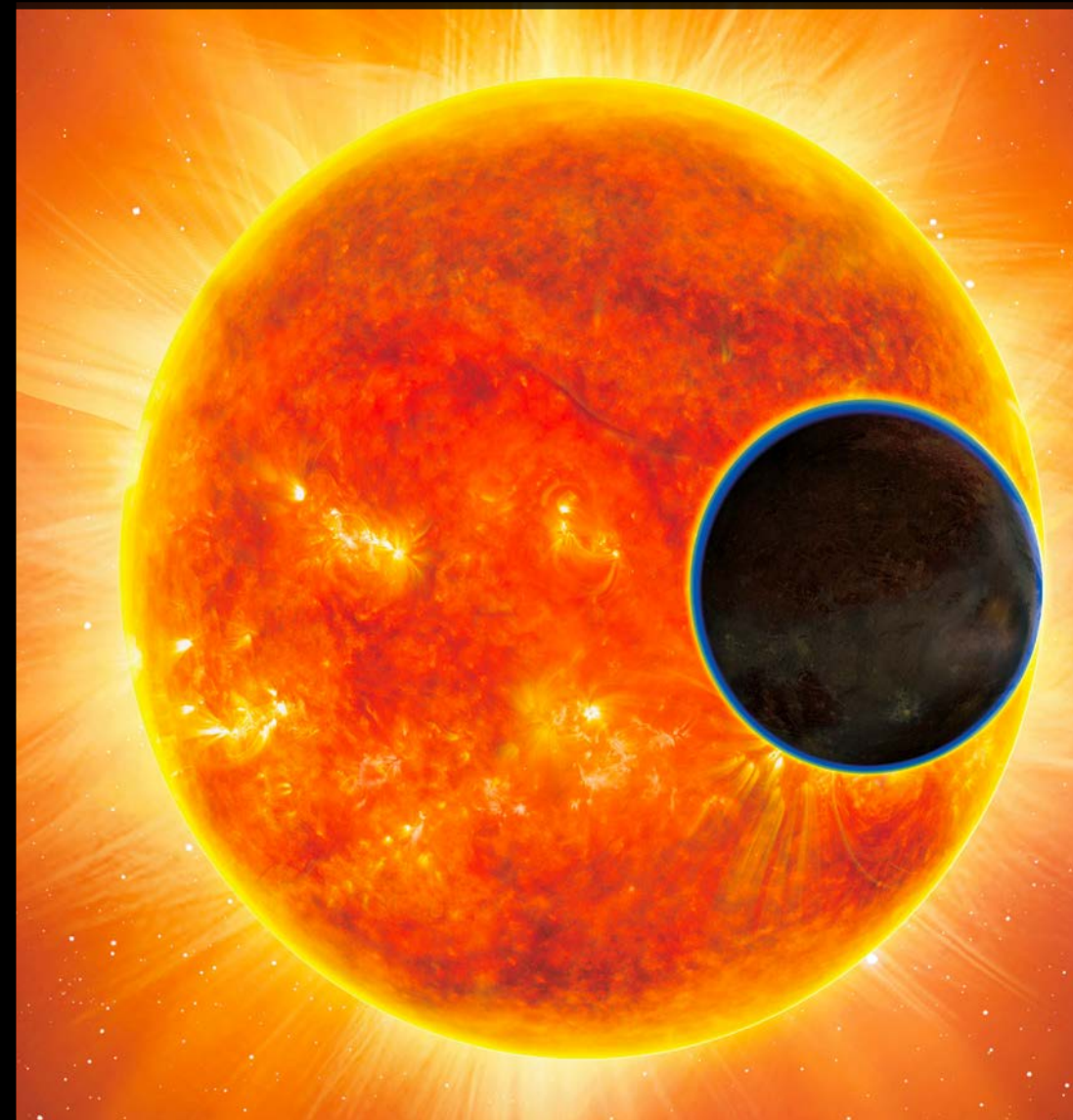
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In situ  
(we go there, boldly)

$N_{\text{stars}} = 1$

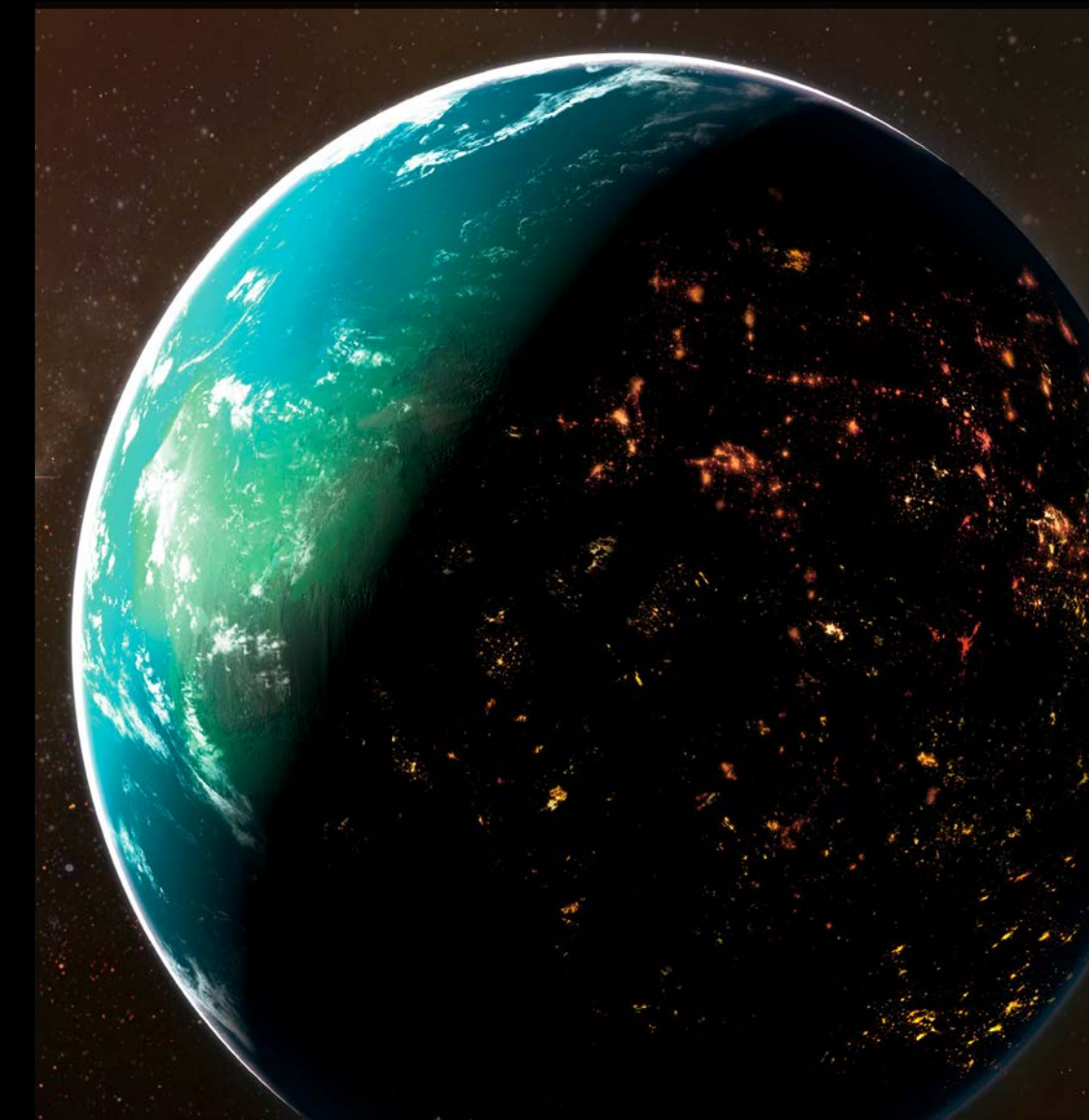
\$\$\$\$\$



Atmospheric biosignature  
(chemical disequilibrium)

$N_{\text{stars}} \sim 10$

\$



Technosignature detection  
(SETI)

$N_{\text{stars}} \sim 10^{23}$



**BREAKTHROUGH**  
**LISTEN**



# BREAKTHROUGH LISTEN

*“THE APOLLO PROGRAM OF SETI”*  
- E. ENRIQUEZ



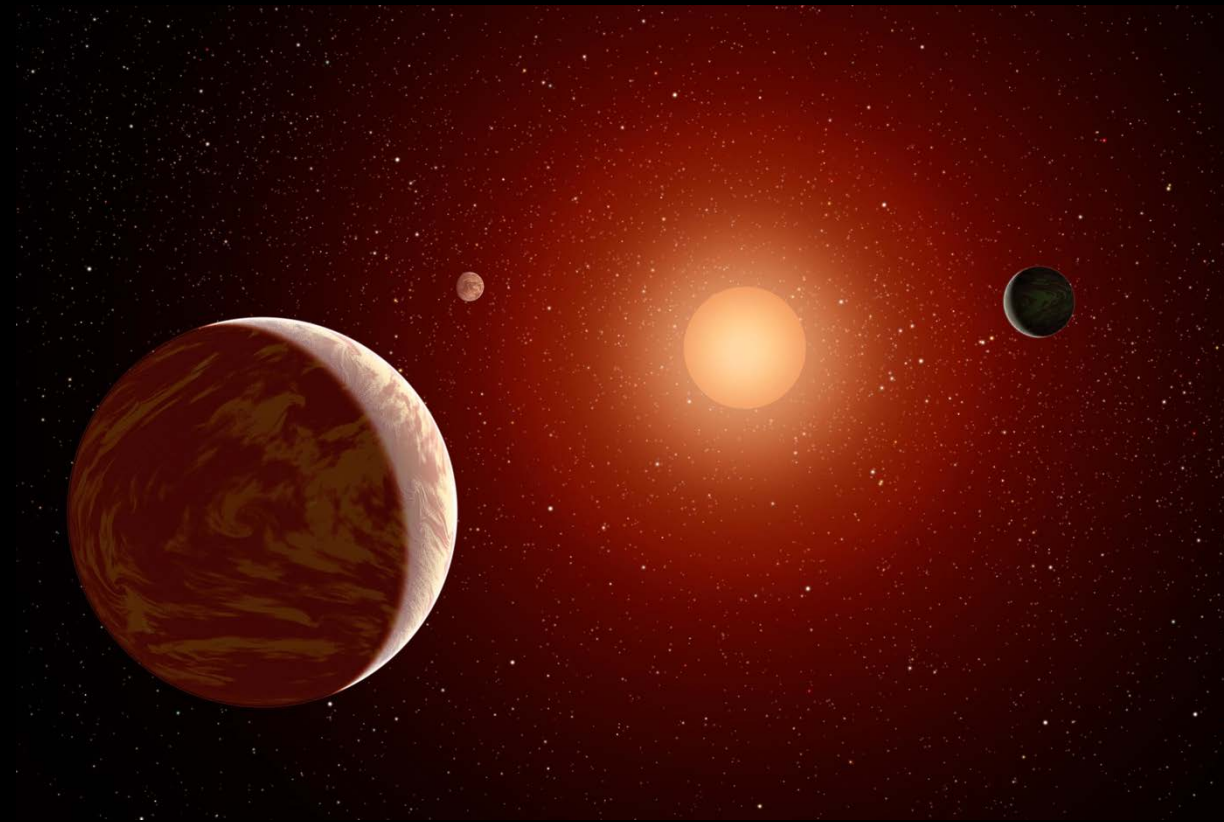


*Breakthrough Listen* is the largest ever scientific research program aimed at finding evidence of intelligent life beyond Earth.



# THE BREAKTHROUGH LISTEN INITIATIVE:

## OVERVIEW



*1 Million Stars*



*MW Galactic Plane Survey*

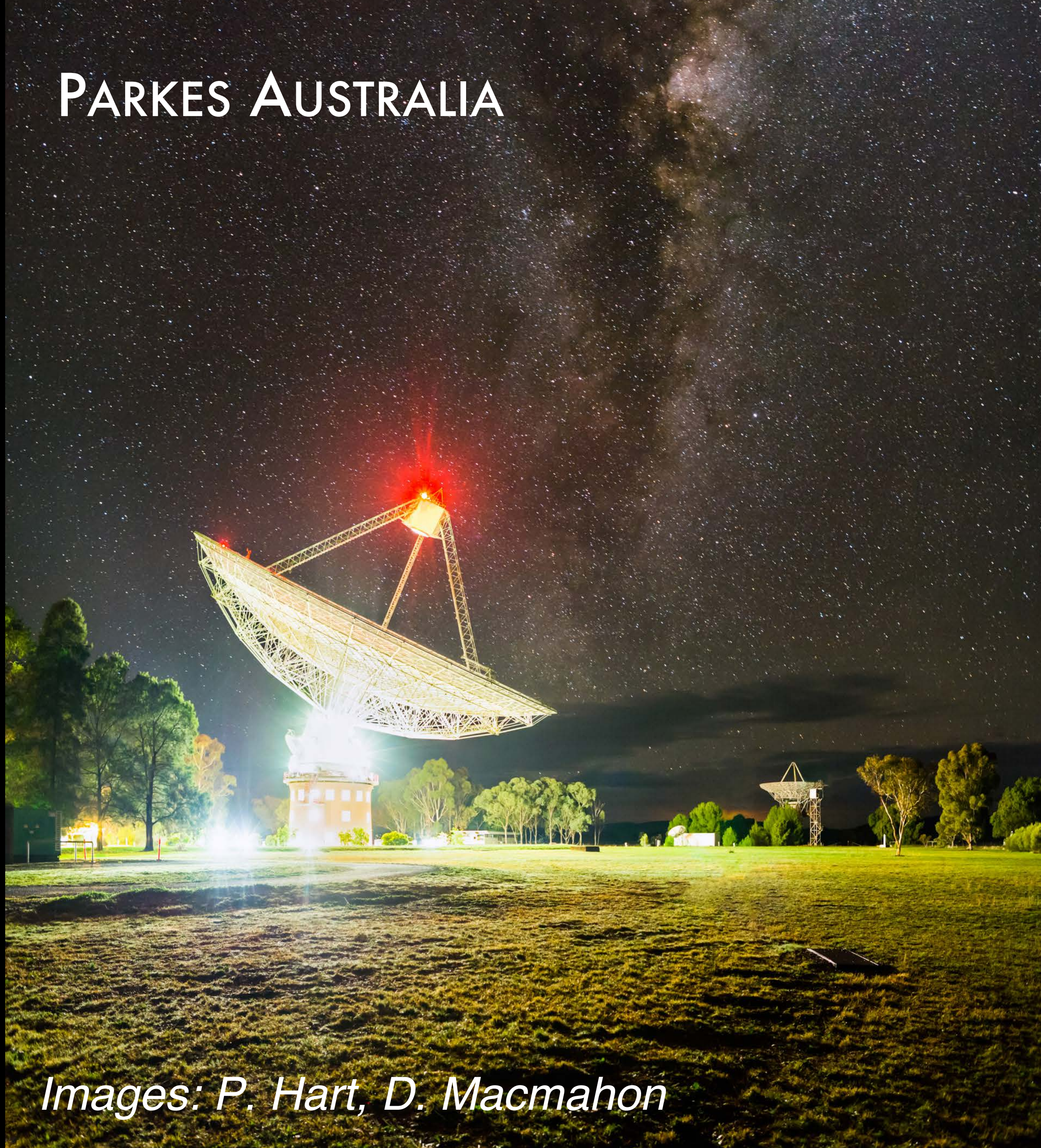


*100 Galaxies*

*Open data and open source*



# PARKES AUSTRALIA



*Images: P. Hart, D. Macmahon*



# GREEN BANK USA



# AUTOMATED PLANET FINDER, LICK OBSERVATORY





# MEERKAT TELESCOPE, SOUTH AFRICA



*Image: SARA0*



# PARTNER FACILITIES

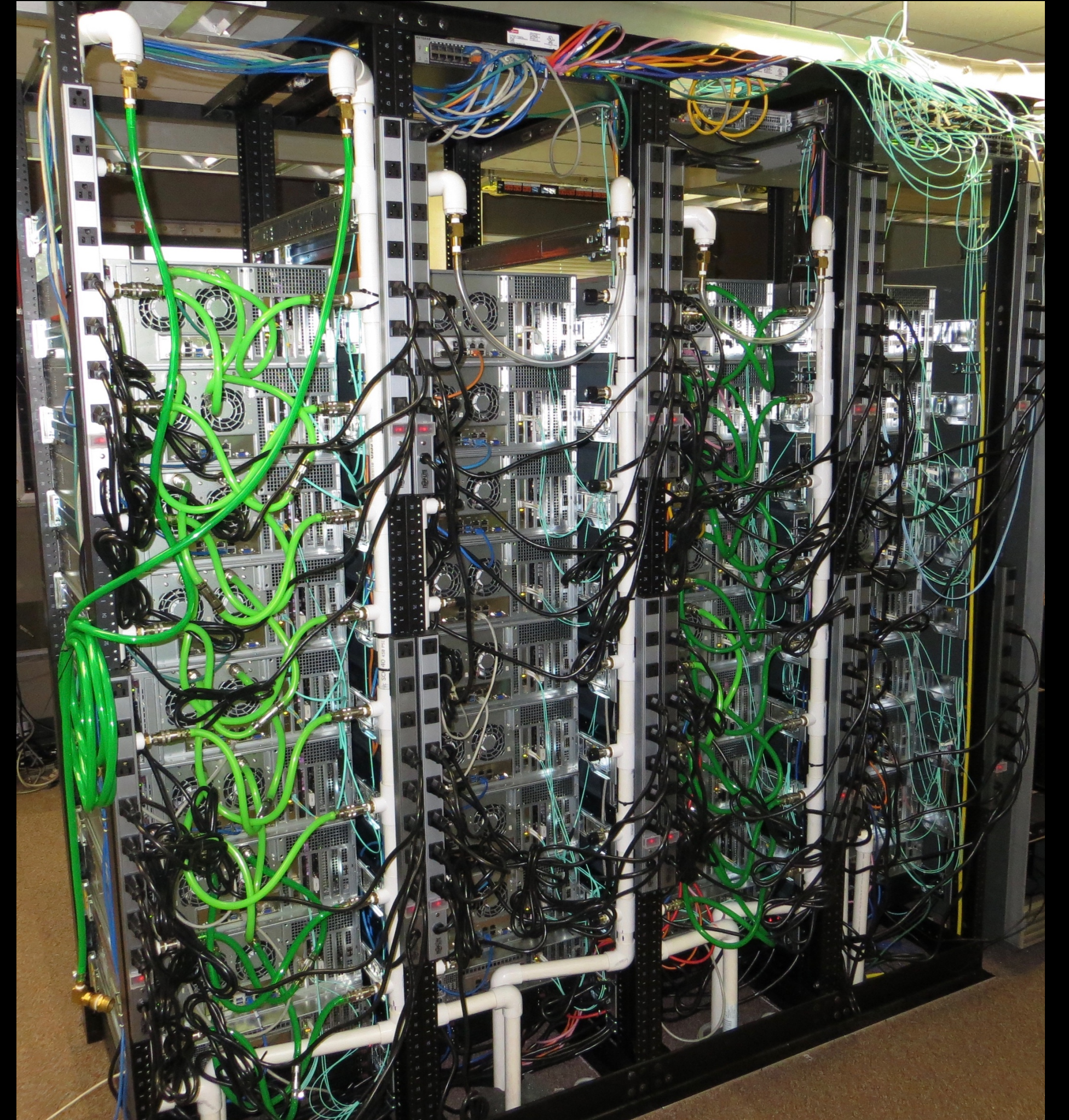




# STEP 1: OBSERVING & RECORDING DATA



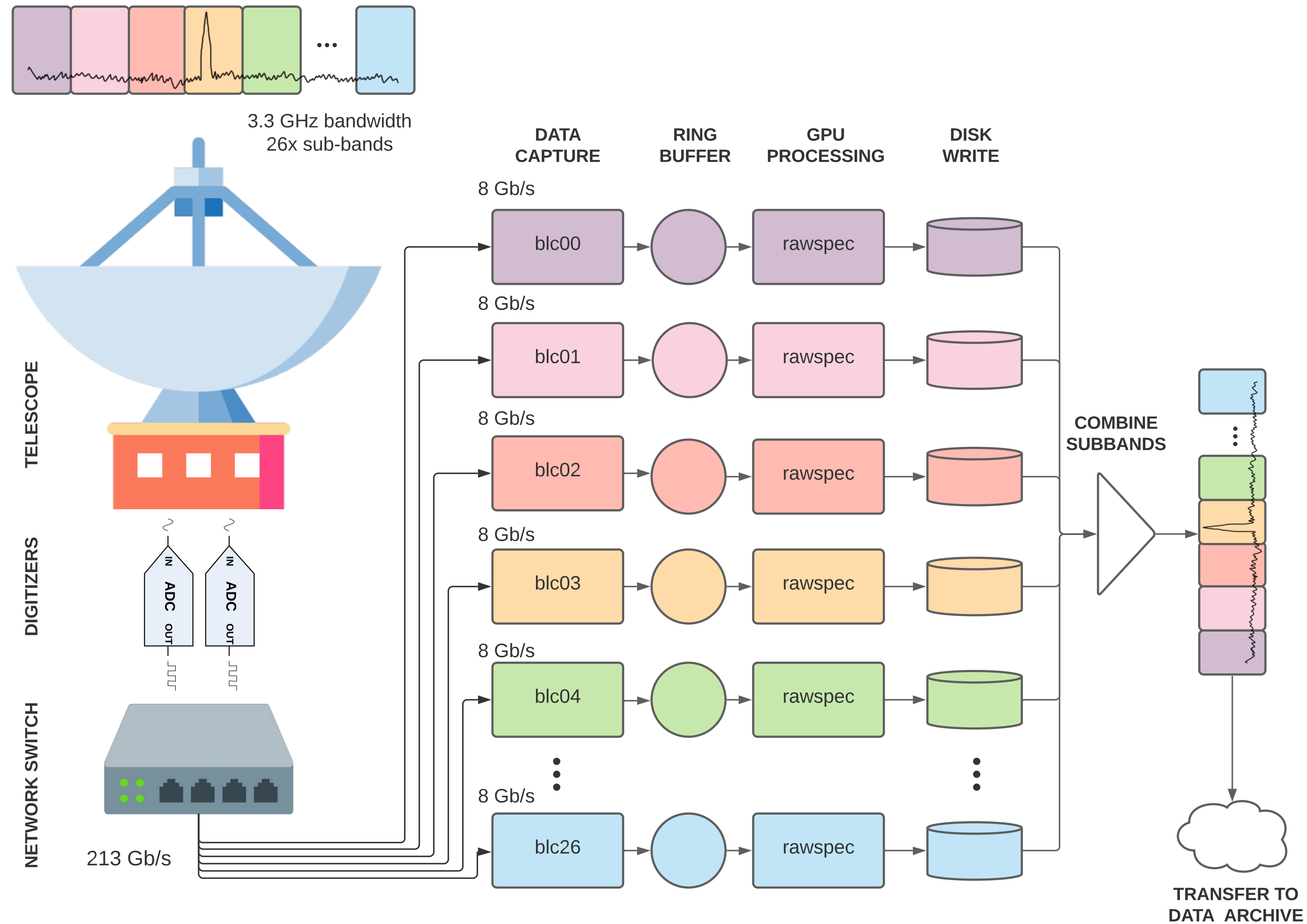
# PARKES AUSTRALIA



# GREEN BANK USA

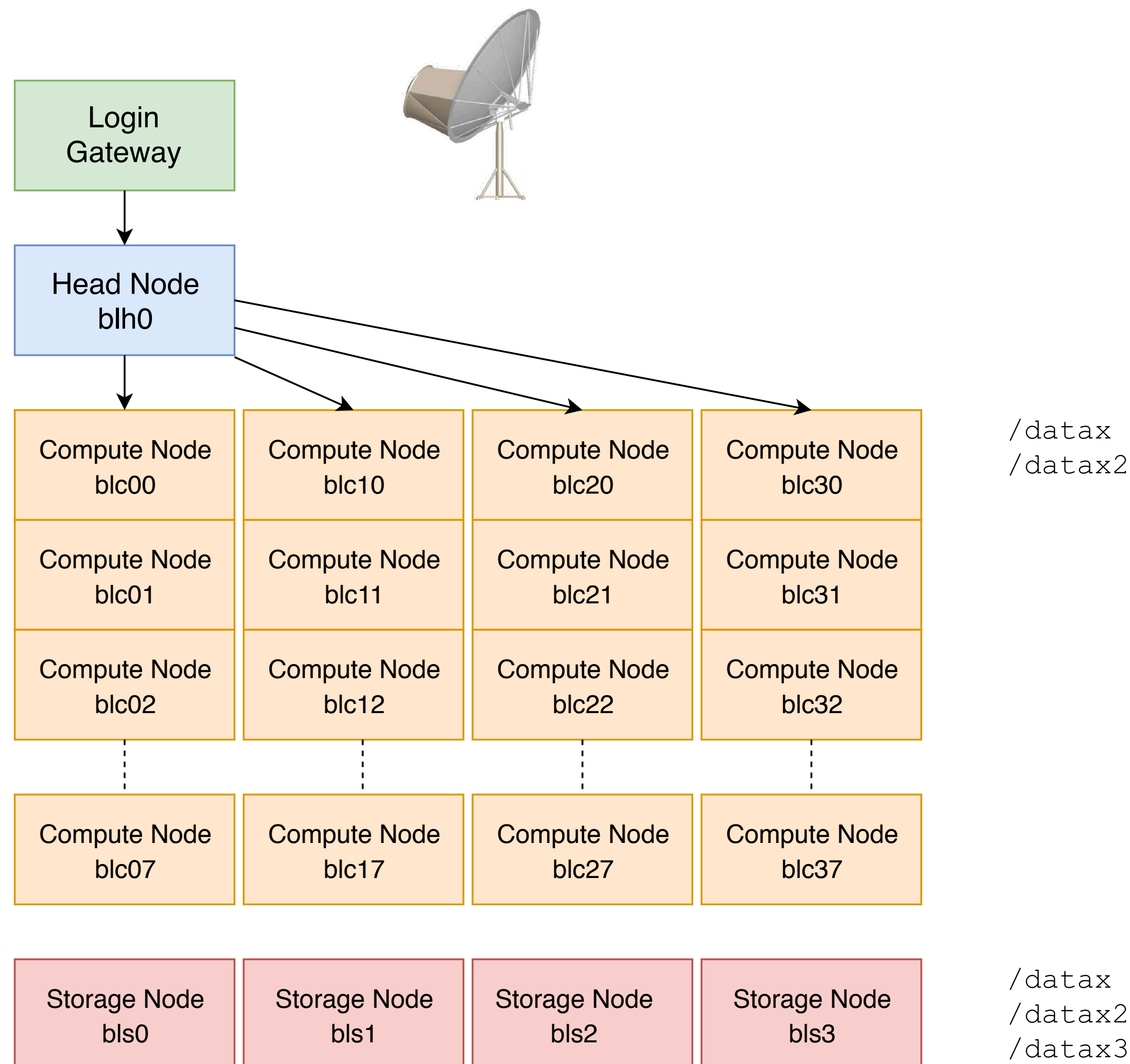


# PARKES DATA FLOW DIAGRAM





# BL DATA RECORDERS



- Currently (Jun 21) 13.6 PB of data stored on disk.
- 9 PB of storage at Green Bank
  - 65x compute nodes (w. GPU)
  - 9x storage nodes (x36 3.5" disks)
- 3.5 PB of storage at Parkes
  - 27x compute nodes
  - 6x storage nodes
- 5 PB of data currently hosted at Berkeley, available at [seti.berkeley.edu/opendata](http://seti.berkeley.edu/opendata)



# DATA CHALLENGES



- High-speed UDP data capture meant we did not use distributed filesystem.
- Started using JBOD (just a bunch of disks), now moving data into a newly-commissioned gluster cluster for archiving.
- The cloud has remained too expensive (about 4x self hosting), but this is improving.
- Started using 4 TB drives (2015), 16 TB drives are now reasonably priced.



## STEP 2: STORING & ANALYZING DATA



# SIGPROC FILTERBANK FORMAT

HEADER

DATA BLOB

- A very simple format with a header followed by a data payload.

```
--- File Info ---
telescope_id :          4
nbits :          32
fch1 :          1361.5
tstart :          58643.2839468
data_type :          1
nchans :          45056
ibeam :          12
tsamp :          0.898779428571
rawdatafile : guppi_58643_24533_053837_G238.12-3.44_0001.0002.fil
foff :          -0.00341796875
src_raj :          7:23:29.904
src_dej :          -24:18:46.8
nbeams :          13
az_start :          0.0
source_name :          G238.12-3.44
za_start :          0.0
machine_id :          20
nifs :          4

Num ints in file :          334
File shape :          (334, 4, 45056)
--- Selection Info ---
Data selection shape :          (334, 4, 45056)
Minimum freq (MHz) :          1207.5
Maximum freq (MHz) :          1361.5
```



# HDF5 FILTERBANK FORMAT

**HDF5 ATTRIBUTES**

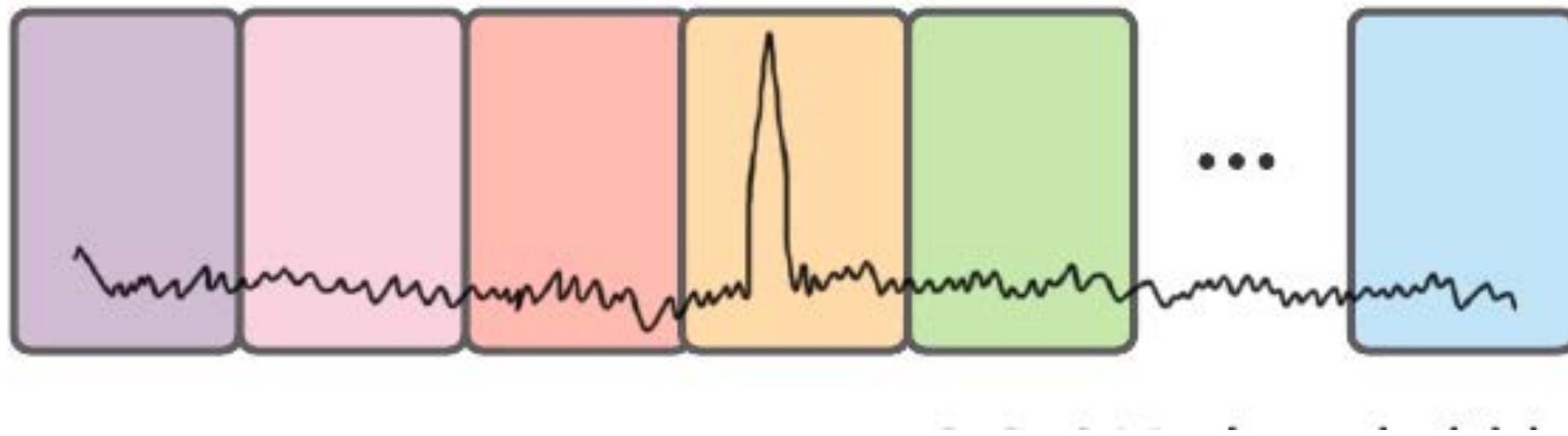
**HDF5 DATASET**

- Filterbank header converted into set of HDF5 attributes.
- Data stored in a HDF5 dataset.
- Applying bitshuffle compression (designed for radio data).
- We use a python package called blimp to interact with sigproc + HDF5 data, uses h5py and hdf5plugin

```
pip install blimp
```



# VIRTUAL DATASET OPPORTUNITIES



- Currently most observations are spread across multiple files (one file per sub-band).
- Can use HDF5 Virtual Datasets to combine sub-bands without moving data.

```
import h5py

nodes          = ['blc0%i'%i for i in range(8)]
n_nodes        = len(nodes)
n_timestep     = 92
n_chan_per_sub = 256
filename       = 'guppi_58948_45245_6051771717_J1019-5749_S_0001.0001.h5x'
vsources = []

layout = h5py.VirtualLayout(shape=(n_timestep, n_chan_per_sub * n_nodes), dtype='<f4')
for ii, node in enumerate(nodes):
    vsource = h5py.VirtualSource(f'collate/{node}/{filename}', 'bp_xx', shape=(92, 256))
    layout[:, ii*n_chan_per_sub:(ii+1)*n_chan_per_sub] = vsource

with h5py.File("VDS_TEST.h5", 'w', libver='latest') as f:
    f.create_virtual_dataset('data', layout, fillvalue=0)
```



## Scaling SETI To The Cloud.



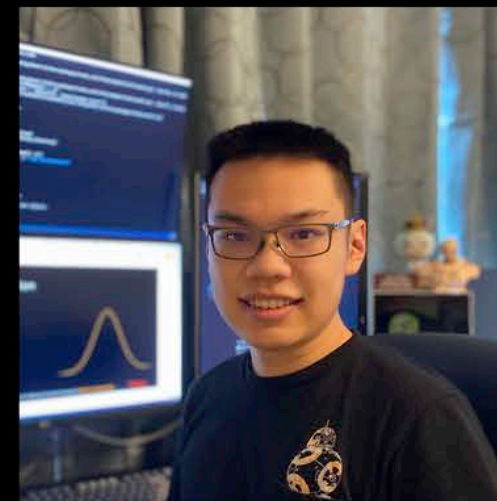
- Our 2020 REU student team created a cloud-based SETI pipeline: BL@Scale.
- Being used to prototype new machine learning algorithms and test cloud workflows.
- We would like to bring these techniques to our on-premises data archive + HPC cluster.



Fatima Zaidouni



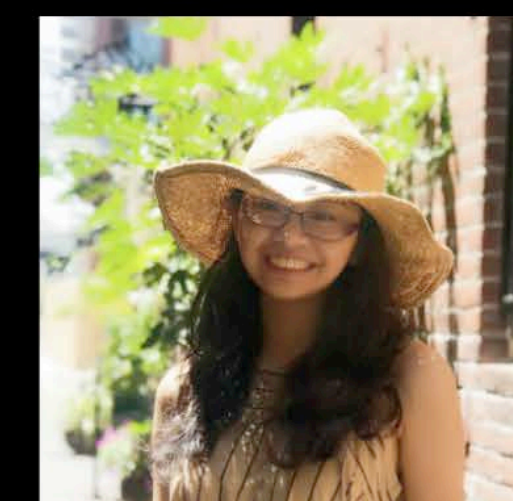
Peter Ma



Yuhong Chen



Shirley Wang



Rachel Zhong



# PROGRESS AND FUTURE STEPS



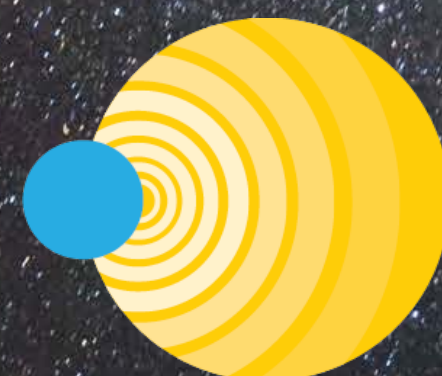
- We have recently rolled out gluster and are expanding our gluster storage capacity.
- We are using Jupyterhub to serve notebooks on co-located GPU servers. We would like to write a SSH spawner to serve notebooks across servers at different sites.
- Can we run HSDS on gluster, and modify our tools to use h5pyd?
- Also considering a SLURM + Singularity processing approach.



# THANK YOU



@berkeleyseti



**BERKELEY SETI**  
RESEARCH CENTER



Image: P. Hart